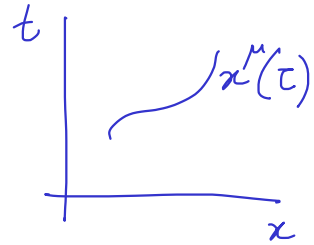


FOUR VECTORS

4-position :

$$x^\mu : \quad x^0 = ct, \quad (x^1, x^2, x^3) = \vec{r}$$

Timelike curves are parametrized by the proper time, τ .



4-velocity :

$$u^\mu = \frac{dx^\mu}{d\tau} = \frac{dx^\mu}{dt} \frac{dt}{d\tau}$$

$$= \gamma \frac{dx^\mu}{dt}$$

$$dt = \gamma d\tau$$

$$\Rightarrow \frac{dt}{d\tau} = \gamma$$

$$= (\gamma c, \gamma \vec{v})$$

4-momentum :

$$p^\mu = m u^\mu = (\gamma mc, \gamma m \vec{v})$$

$m = \text{rest mass}$

$$= (E/c, \vec{p})$$

$$E = \gamma mc^2, \quad \vec{p} = \gamma m \vec{v}$$

4-acceleration :

$$a^\mu = \frac{du^\mu}{d\tau} = \gamma \frac{du^\mu}{dt} = \gamma (c\dot{\gamma}, \dot{\gamma}\vec{v} + \gamma\vec{a}), \quad \dot{\gamma} \equiv \frac{d\gamma}{dt}, \quad \vec{a} \equiv \frac{d\vec{v}}{dt}$$

$$\gamma = \frac{1}{\sqrt{1-\beta^2}} \Rightarrow \dot{\gamma} = \frac{1}{2(1-\beta^2)^{3/2}} \frac{d\beta^2}{dt} = \frac{\gamma^3}{2} \frac{d}{dt} (\vec{v} \cdot \vec{v} / c^2)$$

$$= \frac{\gamma^3}{c^2} \vec{v} \cdot \vec{a}$$

$$\therefore a^\mu = \left(\frac{\gamma^4}{c} \vec{v} \cdot \vec{a}, \frac{\gamma^4}{c^2} (\vec{v} \cdot \vec{a}) \vec{v} + \gamma^2 \vec{a} \right)$$

Proper 4-acceleration (measured in frame where $\vec{v} = 0$)

$$a^\mu_{\text{proper}} = (0, \vec{a})$$

4-force :

$$K^\mu = \frac{dp^\mu}{d\tau} = \gamma \frac{dp^\mu}{dt} = \left(\frac{\gamma}{c} \frac{dE}{dt}, \gamma \frac{d\vec{p}}{dt} \right)$$

$$\vec{F} = \frac{d\vec{p}}{dt}$$

$$\Rightarrow K^\mu = \left(\frac{\gamma}{c} \frac{dE}{dt}, \gamma \vec{F} \right)$$

$$K_{\text{proper}}^\mu = (0, \vec{F}), \quad \vec{v} = 0, \quad E = mc^2 = \text{constant}$$

Scalar product in Minkowski coordinate system

Metric tensor in Minkowski space, $g = \begin{pmatrix} 1 & 0 & 0 & 0 \\ 0 & -1 & 0 & 0 \\ 0 & 0 & -1 & 0 \\ 0 & 0 & 0 & -1 \end{pmatrix}$

$$u \cdot v \equiv u^\alpha v^\beta g_{\alpha\beta} = u^0 v^0 - \vec{u} \cdot \vec{v}$$

$$\Rightarrow u \cdot u = u^\mu u_\mu = u^0 u^0 - \vec{u}^2 \quad \left\{ \begin{array}{l} > 0 \Rightarrow \text{Timelike} \\ < 0 \Rightarrow \text{Spacelike} \\ = 0 \Rightarrow \text{Null / Lightlike} \end{array} \right.$$

$v^\rho g_{\alpha\beta} \equiv v_\alpha \Rightarrow u \cdot v = u^\alpha v_\alpha$

For 4-velocity, $u \cdot u = \gamma^2 c^2 - \gamma^2 \vec{v}^2 = \gamma^2 c^2 (1/\gamma^2) = c^2$
 \therefore The 4-velocity is timelike.

For 4-acceleration, if we look at the proper 4-acceleration,

$$a_{\text{proper}} \cdot a_{\text{proper}} = -\vec{a}^2 \leq 0$$

\Rightarrow The 4-acceleration is spacelike (or null if $\vec{a} = 0$).

For 4-momentum, $p \cdot p = m^2 u \cdot u = m^2 c^2 > 0$ (timelike)

For 4-force, again looking at the proper 4-force,

$$K_{\text{proper}} \cdot K_{\text{proper}} = -\vec{F}^2 \quad (\text{spacelike})$$

$$a \cdot v = \left(\frac{\gamma^4}{c} \vec{v} \cdot \vec{a}, \frac{\gamma^4}{c^2} (\vec{v} \cdot \vec{a}) \vec{v} + \gamma^2 \vec{a} \right) \cdot (\gamma c, \gamma \vec{v})$$

$$= \gamma^5 \vec{v} \cdot \vec{a} - \gamma^5 (\vec{v} \cdot \vec{a}) v^2 / c^2 - \gamma^3 \vec{a} \cdot \vec{v}$$

$$= \gamma^5 \vec{v} \cdot \vec{a} (1/\gamma^2) - \gamma^3 \vec{v} \cdot \vec{a} = 0$$

The scalar product of the 4-velocity and the 4-acceleration is always zero.

Relativistic Kinematics

- Multiparticle collisions

Center of Momentum (COM) frame: Reference frame where the total 3-momentum of all the particles adds up to zero.

Total 4-momentum, $p^\mu = \sum_s p_s^\mu$ s labels particles

$$\begin{aligned} p^\mu p_\mu &= \left(\sum_s \gamma_s m_s c \right)^2 - \left(\sum_s \gamma_s m_s \vec{v}_s \right)^2 \\ &= \sum_{rs} \gamma_r \gamma_s m_r m_s c^2 - \sum_{rs} \gamma_r \gamma_s m_r m_s \vec{v}_r \cdot \vec{v}_s \\ &= \sum_s \underbrace{\gamma_s^2 m_s^2 (c^2 - v_s^2)}_{c^2/\gamma_s^2} + \sum_{r \neq s} \gamma_r \gamma_s m_r m_s (c^2 - \vec{v}_r \cdot \vec{v}_s) \\ &= \sum_s m_s^2 c^2 + \sum_{r \neq s} \gamma_r \gamma_s m_r m_s (c^2 - \vec{v}_r \cdot \vec{v}_s) \geq 0 \quad \star \\ &= \sum_s m_s^2 c^2 + \sum_{r \neq s} p_r^\mu p_{s\mu} \end{aligned}$$

$\star \Rightarrow$ The total 4-momentum is timelike \Rightarrow there exists a frame in which the total 3-momentum vanishes — the COM frame.

Consider a reaction in which two particles collide and produce a set of particles. Consider the 2nd particle to be at rest (stationary target).

$r=1,2 \rightarrow$ Colliding particles; $r=3,4,\dots \rightarrow$ Produced particles

$p^\mu p_\mu$ is invariant in all inertial frames and is conserved in the reaction.

COM frame \rightarrow primed frame.

For initial particles,

In COM frame, $p^{\mu'} = (E/c, 0, 0, 0)$

In the lab frame, $p^\mu p_\mu = (m_1^2 + m_2^2)c^2 + 2p_1^\mu p_{2\mu}$

$$= (m_1^2 + m_2^2)c^2 + 2(E_1 E_2 / c^2 - \vec{p}_1 \cdot \vec{p}_2)$$

$$= (m_1^2 + m_2^2)c^2 + 2m_2 E_1$$

$$= (m_1^2 + m_2^2)c^2 + 2m_2 (T_1 + m_1 c^2)$$

$$= (m_1 + m_2)^2 c^2 + 2m_2 T_1$$

$$\therefore E'^2 / c^2 = (m_1 + m_2)^2 c^2 + 2m_2 T_1$$

$$\Rightarrow E'^2 = (m_1 + m_2)^2 c^4 + 2m_2 c^2 T_1$$

The threshold condition for the reaction corresponds to the situation where the final particles are produced at rest in the COM frame.

At threshold, after the reaction, in the COM frame,

$$p^\mu p_\mu = \left(\sum_r m_r c^2 \right)^2 / c^2 = c^2 \left(\sum_r m_r \right)^2$$

$\sum_r \equiv \sum_{r=3}^{n+2}$ if n particles are produced

$$\therefore (m_1 + m_2)^2 c^2 + 2m_2 T_1^{\text{th}} = c^2 \left(\sum_r m_r \right)^2$$

$$\frac{T_1^{\text{th}}}{m_1 c^2} = \frac{\left(\sum_r m_r \right)^2 - (m_1 + m_2)^2}{2m_1 m_2}$$

Let $Q = \left[\sum_r m_r - (m_1 + m_2) \right] c^2$

$$\therefore \left(\sum_r m_r \right)^2 - (m_1 + m_2)^2 = (Q/c^2) \left[\sum_r m_r + (m_1 + m_2) \right]$$

$$= Q \left[Q/c^2 + 2(m_1 + m_2) \right] / c^2$$

$$\Rightarrow \frac{T_1^{\text{th}}}{m_1 c^2} = \frac{Q^2 + 2Q(m_1 + m_2)c^2}{2m_1 m_2 c^4}$$

Consider the reaction $p+n \rightarrow p+n+p+\bar{p}$

Masses of $p, n, \bar{p} \approx 938 \text{ MeV}/c^2 \equiv m$

\Rightarrow Threshold K.E. = ?

$$Q \approx 2mc^2 \quad \Rightarrow \quad T_1 \approx \frac{4+2 \times 2 \times 2}{2} mc^2 = 6mc^2 = 5.6 \text{ GeV}$$

$$\Rightarrow T_1 = 3Q$$

If we were in the COM frame, all the K.E. would be available for the $p\bar{p}$ production and each incident particle would need a kinetic energy of 938 MeV.

Lagrangian formulation of Special Relativity

[Note: We will only discuss a noncovariant formulation. You can read about covariant formulations on your own.]

A relativistic particle with rest mass m , moving under the influence of a potential energy V , can be described using the Lagrangian,

$$L = -mc^2 \sqrt{1 - \beta^2} - V$$

[In the non relativistic limit, $\sqrt{1 - \beta^2} \sim -\frac{1}{2} \frac{v^2}{c^2} \Rightarrow L = \frac{1}{2} m v^2 - V + \text{constant}$ (as expected).]

Generalization to many particles: $L = -\sum_s m_s c^2 \sqrt{1 - \beta_s^2} - V.$

Equation of motion:

$$\frac{d}{dt} \left(\frac{\partial L}{\partial v_i} \right) = \frac{\partial L}{\partial x_i}$$

$$\frac{\partial L}{\partial v_i} = \frac{-mc^2}{2\sqrt{1-\beta^2}} \left(\frac{-2v_i}{c^2} \right) = \gamma m v_i = p_i, \quad \frac{\partial L}{\partial x_i} = -\frac{\partial V}{\partial x_i} = F_i$$

$$\Rightarrow \frac{d\vec{p}}{dt} = \vec{F}$$

If L doesn't depend on time explicitly, then, the energy is conserved.

$\Rightarrow h = p\dot{q} - L$, considering a single particle in one dimension.

$$= \gamma m v^2 + mc^2 \sqrt{1 - \beta^2} + V$$

$$= \gamma mc^2 \left(\frac{v^2}{c^2} + 1 - \beta^2 \right) + V$$

$$= \gamma mc^2 + V = \text{Total energy.}$$

One dimensional motion under the influence of a constant force

e.g. under the influence of earth's gravity near the earth's surface.

Let the force per unit mass be 'a' $\Rightarrow V = -max$

Let at $t=0$, $x=0$ and $\dot{x}=0$.

[Non-relativistic answer: $x = \frac{1}{2}at^2$ (parabola in $x-t$ plane).]

Using $L = -mc^2\sqrt{1-\beta^2} + max$, we get,

$$\frac{d}{dt}\left(\frac{\partial L}{\partial v}\right) = \frac{\partial L}{\partial x} \Rightarrow \frac{d}{dt}(\gamma mv) = ma$$

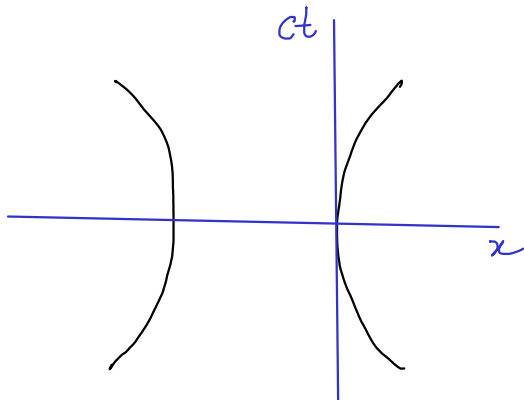
$$\Rightarrow \gamma v = at \quad \because \text{at } t=0, v=0$$

$$\Rightarrow 1-\beta^2 = \left(\frac{v}{at}\right)^2 \Rightarrow v^2\left(\frac{1}{(at)^2} + \frac{1}{c^2}\right) = 1$$

$$\Rightarrow \frac{dx}{dt} = \frac{atc}{\sqrt{(at)^2 + c^2}} = \frac{ct}{\sqrt{t^2 + (c/a)^2}} \quad \text{---(1)}$$

$$\Rightarrow x = c \sqrt{t^2 + (c/a)^2} \Big|_0^t = c \sqrt{t^2 + (c/a)^2} - c^2/a$$

$$\Rightarrow \left(x + c^2/a\right)^2 - c^2t^2 = c^4/a^2 \quad \rightarrow \text{Hyperbola}$$



[To get the non relativistic answer, assume $c^2 \gg at^2$ in Eq. (1)]

$$\Rightarrow \frac{dx}{dt} = at \Rightarrow x = \frac{1}{2}at^2$$